## UNIVERSITY OF CAMBRIDGE INTERNATIONAL EXAMINATIONS GCE Advanced Subsidiary Level and GCE Advanced Level

## MARK SCHEME for the October/November 2010 question paper

## for the guidance of teachers

## 9702 PHYSICS

9702/42 Paper 4 (A2 Structured Questions), maximum raw mark 100

This mark scheme is published as an aid to teachers and candidates, to indicate the requirements of the examination. It shows the basis on which Examiners were instructed to award marks. It does not indicate the details of the discussions that took place at an Examiners' meeting before marking began, which would have considered the acceptability of alternative answers.

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_		www.dynamicpape								
	Page 2	2		ark Scheme: Tea			Syllabus	Paper	•	
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	Section A									
1	(a) for	ce pe	r unit mass	(ratio idea e	essential)			B1	[1]	
	<b>(b)</b> gra		correct curve from ( <i>R</i> ,1.0g	ature g <sub>s</sub> ) & at least one	other correct po	int		M1 A1	[2]	
	(c) (i)			nd Moon are in op	-			M1		
		<i>or</i> so tl	any other	field found by sul r sensible comme nt where it is zero for 2 marks)	ent	eld strength		A1 A0	[2]	
	(ii)		<sub>E</sub> / x <sup>2</sup> = GM <sub>M</sub> / × 10 <sup>24</sup> ) / (7.4 54 R <sub>E</sub>	$((D-x)^2)^4 \times 10^{22} = x^2 / (60)^2$	$(R_{\rm E}-x)^2$			C1 C1 A1	[3]	
	(iii)	(iii) graph: $g = 0$ at least $\frac{2}{3}$ distance to Moon $g_E$ and $g_M$ in opposite directions correct curvature (by eye) and $g_E > g_M$ at surface						B1 M1 A1	[3]	
2	(a) (i)	no f	orces (of attr	action or repulsio	n) between atom	ns / molecule	es / particles	B1	[1]	
	(ii)		of kinetic ar to random m	nd potential energ notion	y of atoms / mole	ecules		M1 A1	[2]	
	(iii)	•	dom) kinetic otential ener	energy increases	s with temperatur	e		M1		
				emperature increa	ases internal ene	ergy)		A1	[2]	
	(b) (i)	zerc	)					A1	[1]	
	(ii)							C1		
								A1	[2]	
	(iii)									
	()		change	work done / J	heating / J	increase i enerç				
			$\begin{array}{c} P \rightarrow Q \\ Q \rightarrow R \end{array}$	<b>+240</b> 0	600 +720	-3 +7				

$Q \rightarrow R$	0	+720	+720
$R \rightarrow P$	<b>-840</b>	+480	-360
(correct signs es	sential)		

(correct signs essential) (each horizontal line correct, 1 mark – max 3)

B3 [3]

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3	(a)	(i)	resona	ance		B1	[1]				
		(ii)	amplit	ude 16mm <u>and</u> frequency 4.6Hz		A1	[1]				
	(b)	(i)	a =4	$-)\omega^2 x$ and $\omega = 2\pi f$ $\pi^2 \times 4.6^2 \times 16 \times 10^{-3}$ $3.4 \mathrm{m  s^{-2}}$		C1 C1 A1	[3]				
		(ii)	F = n = 1	<i>na</i> 50 × 10 <sup>−3</sup> × 13.4		C1					
			= 2	.0N		A1	[2]				
	(c)		•	s 'below' given line and never zero 4.6 Hz (or slightly less) and flatter		M1 A1	[2]				
4	(a)	cha	irge / po	otential (difference) (ratio must be clear)		B1	[1]				
	(b)	(i)	V = Q	$/ 4\pi \varepsilon_0 r$		B1	[1]				
		(ii)	C = Q so C ∘	$V = 4\pi \varepsilon_0 r$ and $\frac{4\pi \varepsilon_0}{10}$ is constant		M1 A0	[1]				
	(c)	(i)	r = C / r = (6. = 6.1 =	$^{\prime}4\pi \varepsilon_{0}r$ 8 × 10 <sup>-12</sup> ) / (4 $\pi$ × 8.85 × 10 <sup>-12</sup> ) × 10 <sup>-2</sup> m		C1 C1 A1	[3]				
		(ii)		$V = 6.8 \times 10^{-12} \times 220$ $1.5 \times 10^{-9}$ C		A1	[1]				
	(d)	(i)	V = Q = 83 V	$C = (1.5 \times 10^{-9}) / (18 \times 10^{-12})$		A1	[1]				
		(ii)	either	energy = $\frac{1}{2}CV^2$ $\Delta E = \frac{1}{2} \times 6.8 \times 10^{-12} \times 220^2 - \frac{1}{2} \times 18 \times 10^{-12} \times 83^2$ = 1.65 × 10 <sup>-7</sup> - 6.2 × 10 <sup>-8</sup>		C1 C1					
			or	= $1.65 \times 10^{-7} - 6.2 \times 10^{-6}$ = $1.03 \times 10^{-7}$ J energy = $\frac{1}{2}$ QV $\Delta E = \frac{1}{2} \times 1.5 \times 10^{-9} \times 220 - \frac{1}{2} \times 1.5 \times 10^{-9} \times 83$ = $1.03 \times 10^{-7}$ J		A1 (C1) (C1) (A1)	[3]				

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	Ра	age 4		Mark Scheme: Teachers' version Syllabus			
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5	(a)	fiel	d into	(the plane of) the paper		B1	[1]
	(b)		² / r = = (2	e to magnetic field <u>provides</u> the centripetal force <i>Bqv</i> 0 × 1.66 × $10^{-27}$ × 1.40 × $10^{5}$ ) / (1.6 × $10^{-19}$ × 6.4 × $10^{-2}$ 454 T	)	B1 C1 B1 A0	[3]
	(c)	(i)	<u>sem</u>	icircle with diameter greater than 12.8 cm		B1	[1]
		(ii)	new	flux density = $\frac{22}{20} \times 0.454$ B = 0.499 T		C1 A1	[2]
							[-]
6	(a)	(i)	e.g.	prevent flux losses / improve flux linkage		B1	[1]
		(ii)	e.m	in core is changing .f. / current (induced) <u>in core</u> iced current in core causes heating		B1 B1 B1	[3]
	(b)	(i)		value of the direct current producing same (mean) pow resistor	er / heating	M1 A1	[2]
		(ii)	-	ver in primary = power in secondary $I_P = V_S I_S$		M1 A1	[2]
7	(a)	(i)	e.g.	electron / particle diffraction		B1	[1]
		(ii)	e.g.	photoelectric effect		B1	[1]
	(b)	(i)	6			A1	[1]
		(ii)	$\lambda = 1$	nge in energy = $4.57 \times 10^{-19}$ J hc / E		C1	
				$.63 \times 10^{-34} \times 3.0 \times 10^{8})$ / (4.57 $\times 10^{-19}$ ) 4 $\times 10^{-7}$ m		A1	[2]
8	(a)			of a heavy nucleus ( <i>not atom/nuclide</i> ) (lighter) nuclei of <u>approximately same mass</u>		M1 A1	[2]
	(b)	⁰n 42H 73Li		(allow $\frac{4}{2}\alpha$ )		M2 A1	[3]
	(c)			particles have kinetic energy	in rods /	B1	
		range of particles in the control rods is short / particles stopped in rods / lose kinetic energy in rods kinetic energy of particles converted to thermal energy				B1 B1	[3]

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				Section B					
9	(a)	(i)	non-	inverting (amplifier)		B1	[1]		
		(ii)	(G =	$(1 + R_2 / R_1)$		B1	[1]		
	(b)	(i)	•	= 1 + 100 / 820 ut = 17 mV		C1 A1	[2]		
		(ii)	( <i>R</i> <sub>2</sub> / (1 +	$R_1$ scores 0 in <b>(a)(ii)</b> but possible 1 mark in each of <b>(b)</b> $R_1 / R_2$ ) scores 0 in <b>(a)(ii)</b> , no mark in <b>(b)(i)</b> , possible 1 i $R_2 / R_1$ ) or $R_1 / R_2$ scores 0 in <b>(a)(ii)</b> , <b>(b)(i)</b> and <b>(b)(ii)</b> )		A1	[1]		
10	(a)	(i)	dens	sity × <u>speed of wave</u> (in the medium)		B1	[1]		
		(ii)	ρ = =	$(7.0 \times 10^{6}) / 4100$ 1700 kg m <sup>-3</sup>		A1	[1]		
	(b)	(i)	I = 1	$T_{T} + I_{R}$		B1	[1]		
		(ii)	<b>1.</b> α	$= (0.1 \times 10^{6})^{2} / (3.1 \times 10^{6})^{2}$ = 0.001		C1 A1	[2]		
			<b>2.</b> α	≈ 1		A1	[1]		
	(c)	eith or		very little transmission at an air-skin boundary (almost) complete transmission at a gel-skin boundary when wave travels in or out of the body no gel, majority reflection with gel, little reflection when wave travels in or out of the body		M1 M1 (M1) (M1) (A1)	[3]		
11	(a)	(i)	unwa	anted random power / signal / energy		B1	[1]		
		(ii)	loss	of (signal) power / energy		B1	[1]		
	(b)	(i)	eithe	er signal-to-noise ratio at mic. = $10 \log (P_2 / P_1)$ = $10 \log (\{2.9 \times 10^{-6}\} / \{3$	.4 × 10 <sup>-9</sup> })	C1			
				= 29  dB maximum length = (29 – 24) / 12 = 0.42 km = 420 m	, , , , , , , , , , , , , , , , , , ,	A1 C1 A1	[4]		
			or	signal-to-noise ratio at receiver = 10 lg $(P_2 / P_1)$ at receiver, 24 = 10 lg $(P / \{3.4 \times 10^{-9}\})$ $P = 8.54 \times 10^{-7}$ W power loss in cables = 10 lg $(\{2.9 \times 10^{-6}\} / \{8.54 \times 10^{-6}\})$ = 5.3 dB	<sup>-7</sup> })	(C1) (A1) (C1)			
				length = 5.3 / 12 km = 440 m		(A1)			

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			d to the m	icrophone ers scores no mark)		M1 A1	[2]
12	(a)	(carrier wave) transmitted from Earth to satellite(1)satellite receives greatly attenuated signal(1)signal amplified and transmitted back to Earth(1)at a different (carrier) frequency(1)different frequencies prevent swamping of uplink signal(1)e.g. of frequencies used (6/4 GHz, 14/11 GHz, 30/20 GHz)(1)(two B1 marks plus any two other for additional physics)(1)		(1) (1)	B1 B1 B2	[4]	
	(b)	advantage:	e.g.	much shorter time delay because orbits are much lower whole Earth may be covered in several orbits / with network		M1 A1 (M1) (A1)	
		disadvanta	ge: e.g.	<i>either</i> must be tracked <i>or</i> limited use in any one orbit more satellites required for continuous of	operation	M1 A1	[4]